

PHY 712 Electrodynamics
12-12:50 AM MWF via video link:
<https://wakeforest-university.zoom.us/my/natalie.holzwarth>

Plan for Lecture 24:

- 1. Brief comment on pure time harmonic radiation from Chap. 9 and HW 18**
- 2. Start reading Chap. 11**
 - A. Equations in cgs (Gaussian) units**
 - B. Special theory of relativity**
 - C. Lorentz transformation relations**

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

1

In this lecture we will jump to Chapter 11 of Jackson and the special theory of relativity. Before we do, there is one point from radiation theory that I would like to clarify.

21	Mon: 03/23/2020	Chap. 9	Radiation from localized oscillating sources	#17	03/25/2020
22	Wed: 03/25/2020	Chap. 9	Radiation from oscillating sources	#18	03/27/2020
23	Fri: 03/27/2020	Chap. 9 and 10	Radiation from oscillating sources	#19	03/30/2020
24	Mon: 03/30/2020	Chap. 11	Special Theory of Relativity	#20	04/03/2020
25	Wed: 04/01/2020	Chap. 11	Special Theory of Relativity		
26	Fri: 04/03/2020	Chap. 11	Special Theory of Relativity		
27	Mon: 04/06/2020	Chap. 14	Radiation from accelerating charged particles		
28	Wed: 04/08/2020	Chap. 14	Synchrotron radiation		
	Fri: 04/10/2020	No class	<i>Good Friday</i>		
29	Mon: 04/13/2020	Chap. 14	Synchrotron radiation		
30	Wed: 04/15/2020	Chap. 15	Radiation from collisions of charged particles		
31	Fri: 04/17/2020	Chap. 13	Cherenkov radiation		
32	Mon: 04/20/2020		Special topic: E & M aspects of superconductivity		
33	Wed: 04/22/2020		Special topic: Aspects of optical properties of materials		
34	Fri: 04/24/2020				
35	Mon: 04/27/2020				
36	Wed: 04/29/2020		Review		

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

2

Here is the undated schedule. Notice that homework #20 is due on Friday this time.

Online colloquium scheduled for Wednesday, April 1, 2020 --
<https://www.physics.wfu.edu/events/colloquium-microstructure-control-in-organic-and-hybrid-semiconductors-and-its-impact-on-device-performance>

Online Colloquium: “Microstructure Control in Organic
and Hybrid Semiconductors and its Impact on Device
Performance “

Public talk for Ph. D. defense

Mr. Andrew Zeidell, Graduate Student

Mentor: Professor Oana Jurchescu

Department of Physics

Wake Forest University

Wednesday, April 1, 2020 at 3:00 PM

Video conference link: (available starting at 2:50 PM)

<https://wakeforest-university.zoom.us/j/534312421>

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

3

Calling to your attention the colloquium for Wednesday this week. Andrew Zeidell is completing his Ph. D. thesis under the mentorship of Professor Jurchescu. This will be the public talk portion of this thesis defense. Hopefully when you are in Andrew's shoes, you will be able to give your public talk in person. This is the last scheduled colloquium for the semester. I have sent you email about the remaining requirements for successfully completing PHY 601.

**Brief comment on pure time harmonic radiation from
Chap. 9 and HW 18**

From Lecture 23 --
Exact equations:

For scalar potential (Lorentz gauge, $k \equiv \frac{\omega}{c}$)

$$\tilde{\Phi}(\mathbf{r}, \omega) = \tilde{\Phi}_0(\mathbf{r}, \omega) + \frac{1}{4\pi\epsilon_0} \int d^3r' \frac{e^{ik|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r}-\mathbf{r}'|} \tilde{\rho}(\mathbf{r}', \omega)$$

For vector potential (Lorentz gauge, $k \equiv \frac{\omega}{c}$)

$$\tilde{\mathbf{A}}(\mathbf{r}, \omega) = \tilde{\mathbf{A}}_0(\mathbf{r}, \omega) + \frac{\mu_0}{4\pi} \int d^3r' \frac{e^{ik|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r}-\mathbf{r}'|} \tilde{\mathbf{J}}(\mathbf{r}', \omega)$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

4

At the end of the last lecture in the discussion of the various exact and approximate schemes for calculating radiation for pure time harmonic sources, there were perhaps some lingering questions which I would like to try to clarify using the example of HW 18.

Still exact

Useful expansion :

$$\frac{e^{ik|\mathbf{r}-\mathbf{r}'|}}{4\pi|\mathbf{r}-\mathbf{r}'|} = ik \sum_{lm} j_l(kr_<) h_l(kr_>) Y_{lm}(\hat{\mathbf{r}}) Y_{lm}^*(\hat{\mathbf{r}}')$$

Within this exact expansion we can sometimes make use of the following Asymptotic behavior:

$$x \ll 1 \quad \Rightarrow j_l(x) \approx \frac{(x)^l}{(2l+1)!!}$$

$$x \gg 1 \quad \Rightarrow h_l(x) \approx (-i)^{l+1} \frac{e^{ix}}{x}$$

Note that this approach is well controlled because it uses a convergent orthogonal function expansion.

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

5

Exact expansion discussed in class and various useful expansion for small or large values of the Bessel/Hankel function arguments.

Another approximation; useful, but not as well controlled

Fields from time harmonic source:

$$\tilde{\Phi}(\mathbf{r}, \omega) = \frac{1}{4\pi\epsilon_0} \int d^3r' \frac{e^{ik|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r}-\mathbf{r}'|} \tilde{\rho}(\mathbf{r}', \omega)$$

For $r \gg r'$: $|\mathbf{r}-\mathbf{r}'| \approx r - \hat{\mathbf{r}} \cdot \mathbf{r}' + \dots$

$$\tilde{\Phi}(\mathbf{r}, \omega) \approx \frac{1}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \int d^3r' e^{-ik\hat{\mathbf{r}} \cdot \mathbf{r}'} \tilde{\rho}(\mathbf{r}', \omega)$$

$$\tilde{\mathbf{A}}(\mathbf{r}, \omega) \approx \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int d^3r' e^{-ik\hat{\mathbf{r}} \cdot \mathbf{r}'} \tilde{\mathbf{J}}(\mathbf{r}', \omega)$$

03/30/20120

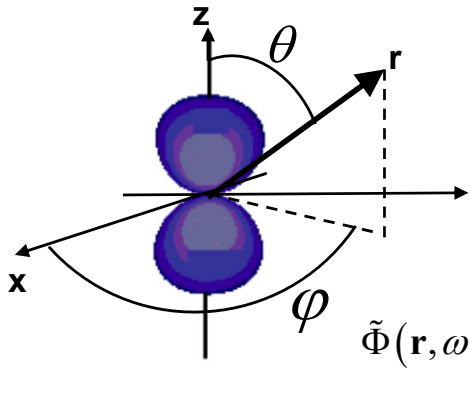
PHY 712 Spring 2020 -- Lecture 24

6

Alternatively, we introduced another approximation scheme which is often called the Born approximation.

For HW 18, you are given the following charge density distribution and asked to evaluate the scalar potential using the exact and dipole approximations. It might be interesting to see what happens when you use the “Born” approximation expression.

Charge density:
$$\rho(\mathbf{r}) = \frac{2e}{\sqrt{6}a_0^4} r \exp\left(-\frac{3r}{2a_0}\right) Y_{00}(\hat{\mathbf{r}})Y_{10}(\hat{\mathbf{r}})e^{-i\omega_0 t}$$



$$Y_{00}(\hat{\mathbf{r}}) = \frac{1}{\sqrt{4\pi}} \quad Y_{10}(\hat{\mathbf{r}}) = \sqrt{\frac{3}{4\pi}} \cos\theta$$

$$\tilde{\Phi}(\mathbf{r}, \omega) \approx \frac{1}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \int d^3r' e^{-ik\hat{\mathbf{r}}\cdot\mathbf{r}'} \tilde{\rho}(\mathbf{r}', \omega)$$

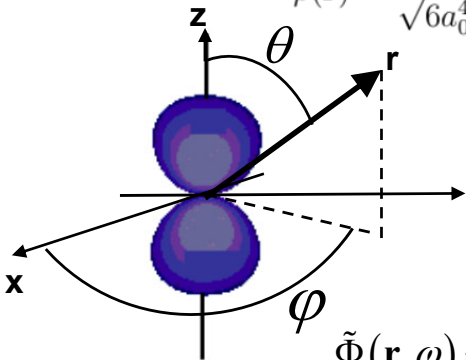
03/30/20120

PHY 712 Spring 2020 -- Lecture 24

7

So the question is how does the Born approximation result differ from the results you used in HW 18? This slide recalls the particular charge distribution of that HW problem.

Charge density: $\rho(\mathbf{r}) = \frac{2e}{\sqrt{6}a_0^4} r \exp\left(-\frac{3r}{2a_0}\right) Y_{00}(\hat{\mathbf{r}})Y_{10}(\hat{\mathbf{r}})e^{-i\omega_0 t}$



$Y_{00}(\hat{\mathbf{r}}) = \frac{1}{\sqrt{4\pi}} \quad Y_{10}(\hat{\mathbf{r}}) = \sqrt{\frac{3}{4\pi}} \cos \theta$

$$\tilde{\Phi}(\mathbf{r}, \omega) \approx \frac{1}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \int d^3r' e^{-ik\hat{\mathbf{r}}\cdot\mathbf{r}'} \tilde{\rho}(\mathbf{r}', \omega)$$

$$\hat{\mathbf{r}} \cdot \mathbf{r}' = r'(\cos \theta \cos \theta' + \sin \theta \sin \theta' \cos(\varphi - \varphi'))$$

$$\int d\Omega' e^{-ik\hat{\mathbf{r}}\cdot\mathbf{r}'} \cos \theta' = -2\pi i j_1(kr' \cos \theta)$$

$$\tilde{\Phi}(\mathbf{r}, \omega) \approx -\frac{i}{\epsilon_0} \frac{e^{ikr}}{r} \frac{Y_{10}(\hat{\mathbf{r}})}{\sqrt{4\pi}} \frac{256\sqrt{6}ka_0e}{(4k^2a_0^2 \cos^2 \theta + 9)^3} \quad \text{different from HW 18}$$

03/30/20120 PHY 712 Spring 2020 -- Lecture 24 8

Using the diagram to evaluate the field position \mathbf{r} and the source position \mathbf{r}' in spherical polar coordinates, we can perform the integrals of the Born approximation. When the dust clears, we can see that the result does differ from the results obtained in HW 18.

Units - SI vs Gaussian

Below is a table comparing SI and Gaussian unit systems. The fundamental units for each system are so labeled and are used to define the derived units.

Variable	SI		Gaussian		SI/Gaussian
	Unit	Relation	Unit	Relation	
length	m	fundamental	cm	fundamental	100
mass	kg	fundamental	gm	fundamental	1000
time	s	fundamental	s	fundamental	1
force	N	$kg \cdot m^2/s$	$dyne$	$gm \cdot cm^2/s$	10^5
current	A	fundamental	$statampere$	$statcoulomb/s$	$\frac{1}{10c}$
charge	C	$A \cdot s$	$statcoulomb$	$\sqrt{dyne \cdot cm^2}$	$\frac{1}{10c}$

03/30/2012

PHY 712 Spring 2020 - Lecture 24

9

Now – jumping to Chapter 11 of Jackson. Fortunately/unfortunately Jackson decided to use cgs Gaussian units starting in Chapter 11. Here is a table of comparison.

Basic equations of electrodynamics

CGS (Gaussian)	SI
$\nabla \cdot \mathbf{D} = 4\pi\rho$	$\nabla \cdot \mathbf{D} = \rho$
$\nabla \cdot \mathbf{B} = 0$	$\nabla \cdot \mathbf{B} = 0$
$\nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}$	$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$
$\nabla \times \mathbf{H} = \frac{4\pi}{c} \mathbf{J} + \frac{1}{c} \frac{\partial \mathbf{D}}{\partial t}$	$\nabla \times \mathbf{H} = \mathbf{J} + \frac{\partial \mathbf{D}}{\partial t}$
$\mathbf{F} = q(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B})$	$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$
$u = \frac{1}{8\pi} (\mathbf{E} \cdot \mathbf{D} + \mathbf{B} \cdot \mathbf{H})$	$u = \frac{1}{2} (\mathbf{E} \cdot \mathbf{D} + \mathbf{B} \cdot \mathbf{H})$
$\mathbf{S} = \frac{c}{4\pi} (\mathbf{E} \times \mathbf{H})$	$\mathbf{S} = (\mathbf{E} \times \mathbf{H})$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

10

More tables of comparison of the two unit schemes.

More relationships

CGS (Gaussian)

$$\mathbf{D} = \mathbf{E} + 4\pi\mathbf{P} = \epsilon\mathbf{E}$$

$$\mathbf{H} = \mathbf{B} - 4\pi\mathbf{M} = \frac{1}{\mu}\mathbf{B}$$

$$\mathbf{E} = -\nabla\Phi - \frac{1}{c}\frac{\partial\mathbf{A}}{\partial t}$$

$$\mathbf{B} = \nabla \times \mathbf{A}$$

MKS (SI)

$$\mathbf{D} = \epsilon_0\mathbf{E} + \mathbf{P} = \epsilon\mathbf{E}$$

$$\mathbf{H} = \frac{1}{\mu_0}\mathbf{B} - \mathbf{M} = \frac{1}{\mu}\mathbf{B}$$

$$\mathbf{E} = -\nabla\Phi - \frac{\partial\mathbf{A}}{\partial t}$$

$$\mathbf{B} = \nabla \times \mathbf{A}$$

$$\epsilon \quad \Leftrightarrow \quad \epsilon / \epsilon_0$$

$$\mu \quad \Leftrightarrow \quad \mu / \mu_0$$

03/30/20120

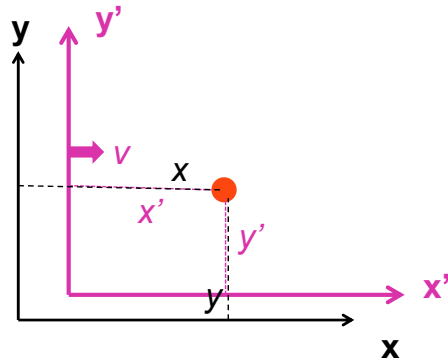
PHY 712 Spring 2020 -- Lecture 24

11

More relationships.

Notions of special relativity

- The basic laws of physics are the same in all frames of reference (at rest or moving at constant velocity).
- The speed of light in vacuum c is the same in all frames of reference.



03/30/20120

PHY 712 Spring 2020 -- Lecture 24

12

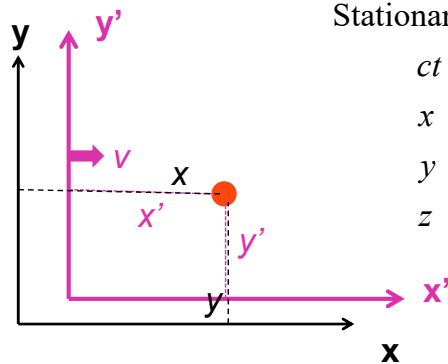
Now – jumping into the story of special relativity. These concepts were covered In Lecture 14 of PHY 742 as well. The black frame corresponds to a (stationary) frame. The purple coordinate system is moving relative to it along the x axis at a speed of v . The red dot is measured differently in the two frames of reference.

Lorentz transformations

Convenient notation :

$$\beta \equiv \frac{v}{c}$$

$$\gamma \equiv \frac{1}{\sqrt{1-\beta^2}}$$



Stationary frame

Moving frame

ct	=	$\gamma(ct' + \beta x')$
x	=	$\gamma(x' + \beta ct')$
y	=	y'
z	=	z'

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

13

The notation with beta and gamma is defined. The question is how the four parameters, ct, x, y, z measured in the stationary reference frame are related to the corresponding four variables ct', x', y', z' measured in the moving frame and vice versa? The consensus is that the Lorentz transformation is the correct correspondence.

Lorentz transformations -- continued

$$\beta \equiv \frac{v}{c} \quad \gamma \equiv \frac{1}{\sqrt{1-\beta^2}}$$

For the moving frame with $\mathbf{v} = v\hat{\mathbf{x}}$:

$$\mathcal{L} = \begin{pmatrix} \gamma & \gamma\beta & 0 & 0 \\ \gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\mathcal{L}^{-1} = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix} = \mathcal{L} \begin{pmatrix} ct' \\ x' \\ y' \\ z' \end{pmatrix}$$

$$\begin{pmatrix} ct' \\ x' \\ y' \\ z' \end{pmatrix} = \mathcal{L}^{-1} \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix}$$

Notice:

$$c^2t^2 - x^2 - y^2 - z^2 = c^2t'^2 - x'^2 - y'^2 - z'^2$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

14

The Lorentz transformation expressed in matrix form. Also note that the four variables have an invariant.

**Examples of other 4-vectors
applicable to the Lorentz transformation:**

For the moving frame with $\mathbf{v} = v\hat{\mathbf{x}}$:

$$\beta \equiv \frac{v}{c} \quad \gamma \equiv \frac{1}{\sqrt{1-\beta^2}}$$

$$\mathcal{L} = \begin{pmatrix} \gamma & \gamma\beta & 0 & 0 \\ \gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\mathcal{L}^{-1} = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\begin{pmatrix} \omega/c \\ k_x \\ k_y \\ k_z \end{pmatrix} = \mathcal{L} \begin{pmatrix} \omega'/c \\ k'_x \\ k'_y \\ k'_z \end{pmatrix}$$

$$\begin{pmatrix} \omega'/c \\ k'_x \\ k'_y \\ k'_z \end{pmatrix} = \mathcal{L}^{-1} \begin{pmatrix} \omega/c \\ k_x \\ k_y \\ k_z \end{pmatrix}$$

Note: $\omega t - \mathbf{k} \cdot \mathbf{r} = \omega' t' - \mathbf{k}' \cdot \mathbf{r}'$
In free space:

$$\left(\frac{\omega}{c}\right)^2 - k^2 = \left(\frac{\omega'}{c}\right)^2 - k'^2 = 0$$

$$\begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix} = \mathcal{L} \begin{pmatrix} E' \\ p'_x c \\ p'_y c \\ p'_z c \end{pmatrix}$$

$$\begin{pmatrix} E' \\ p'_x c \\ p'_y c \\ p'_z c \end{pmatrix} = \mathcal{L}^{-1} \begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix}$$

Note: $E^2 - p^2 c^2 = E'^2 - p'^2 c^2$

Other 4 vectors that obey the Lorentz transformation --

The Doppler Effect

For the moving frame with $\mathbf{v} = v\hat{\mathbf{x}}$:

$$\mathcal{L} = \begin{pmatrix} \gamma & \gamma\beta & 0 & 0 \\ \gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\mathcal{L}^{-1} = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\begin{pmatrix} \omega/c \\ k_x \\ k_y \\ k_z \end{pmatrix} = \mathcal{L} \begin{pmatrix} \omega'/c \\ k'_x \\ k'_y \\ k'_z \end{pmatrix}$$

$$\begin{pmatrix} \omega'/c \\ k'_x \\ k'_y \\ k'_z \end{pmatrix} = \mathcal{L}^{-1} \begin{pmatrix} \omega/c \\ k_x \\ k_y \\ k_z \end{pmatrix}$$

Note: $\omega t - \mathbf{k} \cdot \mathbf{r} = \omega' t' - \mathbf{k}' \cdot \mathbf{r}'$

$$\omega'/c = \gamma(\omega/c - \beta k_x)$$

$$k'_x = \gamma(k_x - \beta\omega/c)$$

$$k'_y = k_y$$

$$k'_z = k_z$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

16

Some details about the frequency/wavevector 4-vector and the Doppler effect for electromagnetic waves.

The Doppler Effect -- continued

$$\omega'/c = \gamma(\omega/c - \beta k_x)$$

$$k'_x = \gamma(k_x - \beta\omega/c)$$

$$k'_y = k_y$$

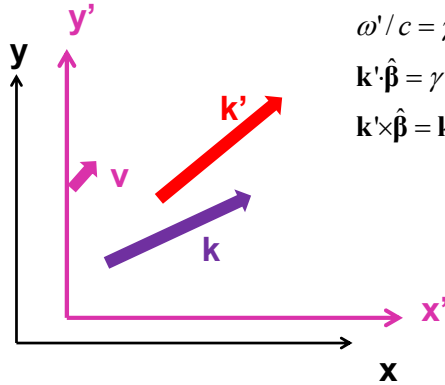
$$k'_z = k_z$$

More generally:

$$\omega'/c = \gamma(\omega/c - \boldsymbol{\beta} \cdot \mathbf{k}) \equiv \gamma(\omega/c - \beta k \cos \theta)$$

$$\mathbf{k}' \cdot \hat{\boldsymbol{\beta}} = \gamma(\mathbf{k} \cdot \hat{\boldsymbol{\beta}} - \beta\omega/c) \equiv k' \cos \theta' = \gamma(k \cos \theta - \beta\omega/c)$$

$$\mathbf{k}' \times \hat{\boldsymbol{\beta}} = \mathbf{k} \times \hat{\boldsymbol{\beta}}$$



For $\theta = 0$: ($k = \omega/c$)

$$\omega' = \omega\gamma(1 - \beta) \quad \Rightarrow \quad \omega' = \omega \sqrt{\frac{1 - \beta}{1 + \beta}}$$

For $\theta \neq 0$: ($k = \omega/c$)

$$\tan \theta' = \frac{\sin \theta}{\gamma(\cos \theta - \beta)}$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

17

Doppler effect continued.

Electromagnetic Doppler Effect ($\theta=0$)

$$\omega' = \omega \sqrt{\frac{1-\beta}{1+\beta}} \quad \beta \approx \frac{v_{\text{source}} - v_{\text{detector}}}{c}$$

More precisely: $\beta = \frac{v_{\text{source}} - v_{\text{detector}}}{c \left(1 - \frac{v_{\text{source}} v_{\text{detector}}}{c^2} \right)}$
(Thanks to E. Carlson)

Sound Doppler Effect ($\theta=0$)

$$\omega' = \omega \left(\frac{1 \pm v_{\text{detector}} / c_s}{1 \mp v_{\text{source}} / c_s} \right)$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

18

Discussion about the Doppler effect for electromagnetic waves and sound waves.

Lorentz transformation of the velocity

Stationary frame Moving frame

$$ct = \gamma(ct' + \beta x')$$

$$x = \gamma(x' + \beta ct')$$

$$y = y'$$

$$z = z'$$

For an infinitesimal increment:

Stationary frame Moving frame

$$cdt = \gamma(cdt' + \beta dx')$$

$$dx = \gamma(dx' + \beta cdt')$$

$$dy = dy'$$

$$dz = dz'$$

Now consider the measurement of velocity in the two different reference frames.

Lorentz transformation of the velocity -- continued

Stationary frame

Moving frame

$$cdt = \gamma(cdt' + \beta dx')$$

$$dx = \gamma(dx' + \beta cdt')$$

$$dy = dy'$$

$$dz = dz'$$

Define:

$$u_x \equiv \frac{dx}{dt} \quad u_y \equiv \frac{dy}{dt} \quad u_z \equiv \frac{dz}{dt}$$

$$u'_x \equiv \frac{dx'}{dt'} \quad u'_y \equiv \frac{dy'}{dt'} \quad u'_z \equiv \frac{dz'}{dt'}$$

$$\frac{dx}{dt} = \frac{\gamma(dx' + \beta cdt')}{\gamma(dt' + \beta dx'/c)} = \frac{u'_x + v}{1 + vu'_x/c^2} = u_x$$

$$\frac{dy}{dt} = \frac{dy'}{\gamma(dt' + \beta dx'/c)} = \frac{u'_y}{\gamma(1 + vu'_x/c^2)} = u_y$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

20

Evaluating the infinitesimals to determine the velocity relationships.

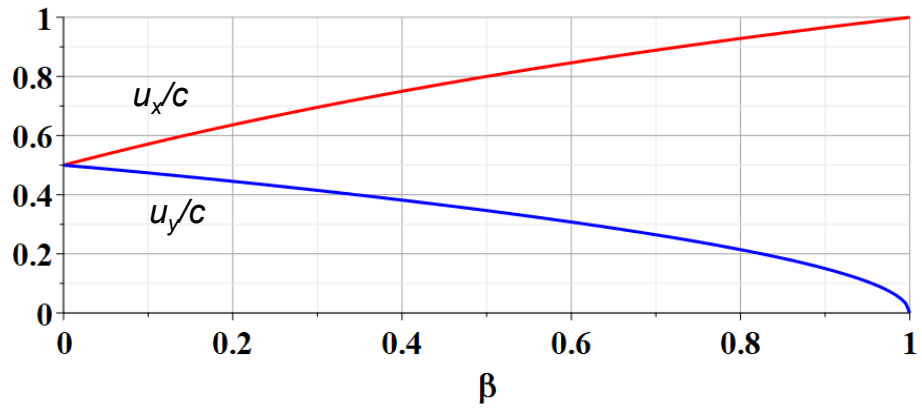
Summary of velocity relationships

$$u_x = \frac{u'_x + v}{1 + vu'_x / c^2}$$

$$u_y = \frac{u'_y}{\gamma(1 + vu'_x / c^2)} \equiv \frac{u'_y}{\gamma_v(1 + vu'_x / c^2)}$$

$$u_z = \frac{u'_z}{\gamma(1 + vu'_x / c^2)} \equiv \frac{u'_z}{\gamma_v(1 + vu'_x / c^2)}$$

Example of velocity variation with β :
($u'_x/c = u'_y/c = 0.5$)



03/30/20120

PHY 712 Spring 2020 -- Lecture 24

22

Numerical evaluation of the velocity relationship.

Extention to transformation of acceleration

$$\mathbf{a}_{\parallel} = \frac{\left(1 - \frac{v^2}{c^2}\right)^{3/2}}{\left(1 + \frac{\mathbf{v} \cdot \mathbf{u}'}{c^2}\right)^3} \mathbf{a}'_{\parallel}$$

$$\mathbf{a}_{\perp} = \frac{\left(1 - \frac{v^2}{c^2}\right)}{\left(1 + \frac{\mathbf{v} \cdot \mathbf{u}'}{c^2}\right)^3} \left(\mathbf{a}'_{\perp} + \frac{\mathbf{v}}{c^2} \times (\mathbf{a}' \times \mathbf{u}') \right)$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

23

It is possible to take the derivatives of the velocities to get the accelerations. The proof of these results are left for you to fill in.

Velocity transformations continued:

Consider: $u_x = \frac{u'_x + v}{1 + vu'_x/c^2}$ $u_y = \frac{u'_y}{\gamma_v(1 + vu'_x/c^2)}$ $u_z = \frac{u'_z}{\gamma_v(1 + vu'_x/c^2)}$.

Note that $\gamma_u \equiv \frac{1}{\sqrt{1 - (u/c)^2}} = \frac{1 + vu'_x/c^2}{\sqrt{1 - (u'/c)^2} \sqrt{1 - (v/c)^2}} = \gamma_v \gamma_{u'} (1 + vu'_x/c^2)$

$\Rightarrow \gamma_u c = \gamma_v (\gamma_{u'} c + \beta_v \gamma_{u'} u'_x)$

$\Rightarrow \gamma_u u_x = \gamma_v (\gamma_{u'} u'_x + \gamma_{u'} v) = \gamma_v (\gamma_{u'} u'_x + \beta_v \gamma_{u'} c)$

$\Rightarrow \gamma_u u_y = \gamma_{u'} u'_y$ $\gamma_u u_z = \gamma_{u'} u'_z$

Velocity 4-vector:
$$\begin{pmatrix} \gamma_u c \\ \gamma_u u_x \\ \gamma_u u_y \\ \gamma_u u_z \end{pmatrix} = \mathcal{L}_v \begin{pmatrix} \gamma_{u'} c \\ \gamma_{u'} u'_x \\ \gamma_{u'} u'_y \\ \gamma_{u'} u'_z \end{pmatrix}$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

24

It is apparent that the velocity 4 vector itself does not obey the Lorentz transformation. These identities show that we can construct a related 4 vector that does obey the Lorentz transformation.

Some details:

$$\gamma_u = \gamma_v \gamma_{u'} \left(1 + v u'_x / c^2\right) \Rightarrow \left(1 - \frac{v^2}{c^2}\right) \left(1 - \frac{u'^2}{c^2}\right) = \left(1 - \frac{u^2}{c^2}\right) \left(1 + \frac{u_x v}{c^2}\right)^2$$

$$\text{where } u_x = \frac{u'_x + v}{1 + v u'_x / c^2} \quad u_y = \frac{u'_y}{\gamma_v (1 + v u'_x / c^2)} \quad u_z = \frac{u'_z}{\gamma_v (1 + v u'_x / c^2)}$$

$$\left(\frac{u_x^2}{c^2} + \frac{u_y^2}{c^2} + \frac{u_z^2}{c^2}\right) \left(1 + \frac{u_x v}{c^2}\right)^2 = \left(\frac{u'_x}{c} + \frac{v}{c}\right)^2 + \left(\frac{u'^2_y}{c^2} + \frac{u'^2_z}{c^2}\right) \left(1 - \frac{v^2}{c^2}\right)$$

$$\frac{u^2}{c^2} \left(1 + \frac{u_x v}{c^2}\right)^2 = \frac{u'^2}{c^2} \left(1 - \frac{v^2}{c^2}\right) + \left(1 + \frac{u_x v}{c^2}\right)^2 - \left(1 - \frac{v^2}{c^2}\right)$$

$$\Rightarrow \left(1 - \frac{u^2}{c^2}\right) \left(1 + \frac{u_x v}{c^2}\right)^2 = \left(1 - \frac{u'^2}{c^2}\right) \left(1 - \frac{v^2}{c^2}\right)$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

25

Some identities that can be proven...

Significance of 4-velocity vector:
$$\begin{pmatrix} \gamma_u c \\ \gamma_u u_x \\ \gamma_u u_y \\ \gamma_u u_z \end{pmatrix}$$

Introduce the “rest” mass m of particle characterized by velocity \mathbf{u} :

$$mc \begin{pmatrix} \gamma_u c \\ \gamma_u u_x \\ \gamma_u u_y \\ \gamma_u u_z \end{pmatrix} = \begin{pmatrix} \gamma_u mc^2 \\ \gamma_u mu_x c \\ \gamma_u mu_y c \\ \gamma_u mu_z c \end{pmatrix} = \begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix}$$

Properties of energy-moment 4-vector:

$$\begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix} = \mathcal{L} \begin{pmatrix} E' \\ p'_x c \\ p'_y c \\ p'_z c \end{pmatrix} \quad \begin{pmatrix} E' \\ p'_x c \\ p'_y c \\ p'_z c \end{pmatrix} = \mathcal{L}^{-1} \begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix} \quad \text{Note: } E^2 - p^2 c^2 = E'^2 - p'^2 c^2$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

26

When the dust clears, the related physical parameters are the energy-momentum 4 vector.

Properties of Energy-momentum 4-vector --
continued

$$\begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix} = \begin{pmatrix} \gamma_u m c^2 \\ \gamma_u m u_x c \\ \gamma_u m u_y c \\ \gamma_u m u_z c \end{pmatrix}$$

Note: $E^2 - p^2 c^2 = \frac{(m c^2)^2}{1 - \beta_u^2} \left(1 - \left(\frac{u_x}{c}\right)^2 - \left(\frac{u_y}{c}\right)^2 - \left(\frac{u_z}{c}\right)^2 \right) = (m c^2)^2 = E'^2 - p'^2 c^2$

Notion of "rest energy": For $\mathbf{p} \equiv 0$, $E = m c^2$

Define kinetic energy: $E_K \equiv E - m c^2 = \sqrt{p^2 c^2 + m^2 c^4} - m c^2$

Non-relativistic limit: If $\frac{p}{m c} \ll 1$, $E_K = m c^2 \left(\sqrt{1 + \left(\frac{p}{m c}\right)^2} - 1 \right)$
 $\approx \frac{p^2}{2m}$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

27

In order to relate the equations to the non-relativistic treatments, we must use the same zero of energy for both. The kinetic energy of a relativistic free particle is related to the energy $E - m c^2$.

Summary of relativistic energy relationships

$$\begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix} = \begin{pmatrix} \gamma_u m c^2 \\ \gamma_u m u_x c \\ \gamma_u m u_y c \\ \gamma_u m u_z c \end{pmatrix}$$

$$E = \sqrt{p^2 c^2 + m^2 c^4} = \gamma_u m c^2$$

$$\text{Check: } \sqrt{p^2 c^2 + m^2 c^4} = m c^2 \sqrt{\gamma_u^2 \beta_u^2 + 1} = \gamma_u m c^2$$

Example: for an electron $m c^2 = 0.5 \text{ MeV}$

for $E = 200 \text{ GeV}$

$$\gamma_u = \frac{E}{m c^2} = 4 \times 10^5$$

$$\beta_u = \sqrt{1 - \frac{1}{\gamma_u^2}} \approx 1 - \frac{1}{2\gamma_u^2} \approx 1 - 3 \times 10^{-12}$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

28

This slide gives some numerical relationships for a highly accelerated electron.

Special theory of relativity and Maxwell's equations

Continuity equation: $\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = 0$

Lorenz gauge condition: $\frac{1}{c} \frac{\partial \Phi}{\partial t} + \nabla \cdot \mathbf{A} = 0$

Potential equations: $\frac{1}{c^2} \frac{\partial^2 \Phi}{\partial t^2} - \nabla^2 \Phi = 4\pi\rho$

$$\frac{1}{c^2} \frac{\partial^2 \mathbf{A}}{\partial t^2} - \nabla^2 \mathbf{A} = \frac{4\pi}{c} \mathbf{J}$$

Field relations: $\mathbf{E} = -\nabla\Phi - \frac{1}{c} \frac{\partial \mathbf{A}}{\partial t}$

$$\mathbf{B} = \nabla \times \mathbf{A}$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

29

All of the previous equations represent relativistic mechanics. Now we want to relate the ideas to electromagnetic theory. We have said that Maxwell's equations already are consistent with the theory of relativity. But we still have some work to do in order to relate the measured fields and sources in two different reference frames. The idea is to guess the correct 4 vectors.

More 4-vectors:

Time and position :

$$\begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix} \Rightarrow x^\alpha$$

Charge and current :

$$\begin{pmatrix} c\rho \\ J_x \\ J_y \\ J_z \end{pmatrix} \Rightarrow J^\alpha$$

Vector and scalar potentials :

$$\begin{pmatrix} \Phi \\ A_x \\ A_y \\ A_z \end{pmatrix} \Rightarrow A^\alpha$$

03/30/20120

PHY 712 Spring 2020 -- Lecture 24

30

Here are our guesses.

Lorentz transformations

$$\mathcal{L}_v = \begin{pmatrix} \gamma_v & \gamma_v \beta_v & 0 & 0 \\ \gamma_v \beta_v & \gamma_v & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Time and space :

$$x^\alpha = \mathcal{L}_v x'^\alpha \equiv \mathcal{L}_v^{\alpha\beta} x'^\beta$$

Charge and current :

$$J^\alpha = \mathcal{L}_v J'^\alpha \equiv \mathcal{L}_v^{\alpha\beta} J'^\beta$$

Vector and scalar potential : $A^\alpha = \mathcal{L}_v A'^\alpha \equiv \mathcal{L}_v^{\alpha\beta} A'^\beta$

These 4 vectors obey the Lorentz transformations. Here we use the notation that repeated indices should be summed over the 4 components. In this case beta is the summed index. Next time we will see how the E and B fields are represented in terms of the Lorentz transformations.